GPM662C User Manual

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1 Description and Features

The GPM662C camera possesses an exceptionally strong capability to capture infrared light (with a high QE value at 650nm). It also has a large full well capacity, high dynamic range, high sensitivity, and low readout noise among other advantages. This results in outstanding performance in planetary photography, as well as in solar and lunar photography.

The features of GPM662C are listed below:

- IMX 662 Color CMOS Sensor
- Resolution: 1920 x 1080
- 2.9 um Square Pixels
- 1/2.8-inch Optical Format
- 12-bit ADC
- Built-in Frame Buffer
- G Sensitivity: 5970mv with 1/30s
- Noise: 0.48 to 6.81e-
- 18 FPS at all Pixel Readout 8 bit
- SNR Max: 45.9 dB
- Dynamic Range: 75.2 dB

2 GPM662C Specifications and Performance

2.1 Camera Specifications

Table 1 GPM662C Technique Specifications.

Sensor	Sony IMX662 back illuminated sens	or			
Diagonal	6.45 mm				
Image Resolution	2.1 mega pixels (1920*1080)				
Pixel Size	2.9μm × 2.9μm				
Image Area	5.57mm × 3.13mm				
M. FRG (P. 17)	12bit	8bit			
Max FPS at Resolution	8.9 FPS @ 1920*1080	17.8 FPS @ 1920*1080			
Shutter Type	Rolling shutter				
Exposure Time	0.1ms - 1000s				
Gain	1x - 150x				
SNR	45.9 dB				
Dynamic Range	75.2 dB				
Read Noise	6.81 – 0.48 e-				
QE Peak	>91%				
Full Well	39.034ke-	39.034ke-			
ADC	12bit	12bit			
Frame Buffer	Built-in	Built-in			
Connection Port	USB2.0 Type C	USB2.0 Type C			
Camera Adaptor	Standard 1.25" for telescope connect	tion, and standard CS adapter for industrial lens			
Protect Windows	IR-cut filter/AR-window				
Spectral Range	380-690nm (with IR Cut window)				
Capture/Control SDK	Windows/Linux/macOS/Android M Python, Java, DirectShow, Twain, etc	Multiple Platform SDK(Native C/C++, C#/VB.NET, e.);			
Recording System	Still picture and movie				
Camera Dimensions	Diameter 37mm * height 72.4mm				
Camera Weight	65.3 gram				
Back Focus Distance	Back Focus Distance is 8.5mm, 17.5	Back Focus Distance is 8.5mm, 17.5mm with C adapter, 12.5 with CS adapter			
Cooling:	Passive cooling	Passive cooling			
Supported OS	Microsoft® Windows® XP / Vista / 7 OSx(Mac OS X) Linux	7 / 8 /10 (32 & 64 bit)			

2.2 Sony IMX662 Sensitivity

The sensor G Sensitivity of GPM662C is 5970mv with 1/30s. Its spectral sensitivity is shown in Figure 1. The sensitivity is measured with a purple excitation LED light source with a color temperature of 2850 K, an IR cut filter CM700(t =1.0 mm) is placed between the LED light source and the sensor receiving surface to irradiate substantially parallel light.

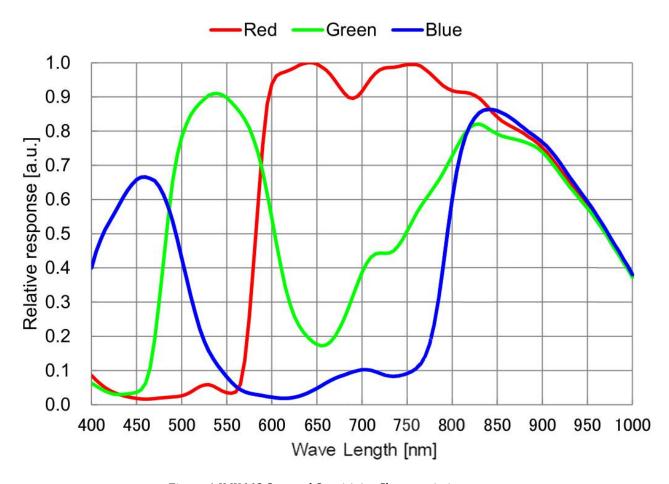


Figure 1 IMX662 Spectral Sensitivity Characteristic

2.3 12bit ADC and ROI

The GPM662C has built in 12bit ADC. It also has 12bit output mode for hardware binning and smaller resolution. The camera also supports hardware ROI, and the smaller the ROI size is, the higher the frame rate is.

Table 2 shows the frame rate of the GPM662C in 12/8bit mode, USB2.0 data transfer interface at different resolutions:

Table 2 GPM662C Frame Rate at Different Resolution/Data Bit/Data Transfer (USB2.0)

Bit & Interface FPS	12bit ADC	8bit ADC
Resolution	USB2.0	USB 2.0
1920*1080	8.9	17.8

2.4 Frame Buffer

GPM662C camera has a built-in frame buffer, which helps maintain the stability of data transmission, and effectively reduce the amp-glow caused because image data can be temporarily buffered without being sent hastily to the receiver. Ensure that each camera can still achieve its maximum frame rate when multiple cameras are operating simultaneously.

2.5 Binning

The GPM662C supports digital binning from 1×1 to 8×8 in either stacking or averaging method.

2.6 Conversion Gain Switch

GPM662C support HCG \ LCG mode switch.

2.7 Power and Cooling System for Precise Temperature Regulation

The camera operates via a USB 2.0 interface. Upon establishing a connection with the host system using the USB 2.0 cable, the device is primed for operation.

2.8 Camera Performance Analysis

Camera performance can be evaluated with e-/ADU, Read Noise, Full Well and Dynamic Range.

e-/ADU: The sensors found in cameras used for vision applications have pixels that convert incoming photons into electrons. Gain on a CCD /CMOS camera represents the conversion factor from electrons (e-) into digital counts, or Analog-Digital Units (ADUs). Gain is expressed as the number of electrons that get converted into a digital number, or electrons per ADU (e-/ADU).

Read Noise: Read Noise is the most important reference to measure the performance of a camera. Lower read noise usually means better SNR and better quality of image. Read Noise is created within the camera electronics during the readout process as the electrons are subjected to the analog to digital conversion, amplification and processing steps that enable an image to be produced.

Full Well: The electrons are held in each pixel and are converted into electrical charge which can be measured to show the amount of light that has fallen on each pixel. The maximum electrical charge possible is termed "full well capacity". Under the same conditions such as noise and A/D converter quality, the greater full well capacity a sensor has, the wider dynamic range the sensor has. As there is a limit to the depth to which pixels can be made, the full well capacity is often proportional to the frontal area of the light gathering element of the pixel.

Dynamic Range is the ratio between the maximum output signal level and the noise floor at minimum signal amplification (noise floor which is the RMS (root mean square) noise level in a black image). The noise floor of the camera contains sensor readout noise, camera processing noise and the dark current shot noise. Dynamic range represents the camera's ability to display/reproduce the brightest and darkest portions of the image and how many variations in between. This is technically intra-scene dynamic range. Within one image there may be a portion that is in complete black and a portion that is completely saturated.

For the GPM series camera, the Gain Value is in xxx% mode. Here xxx is used as the x axis (Gain Value) for the description of the camera performance

$$Rel\ Gain(dB) = 20 * log_{10}[xxx(Gain\ Value)/100]$$

 $xxx(Gain\ Value) = 100 \times 10^{(Rel\ Gain(dB)/20)}$

Read noise is the most important reference to measure the performance of a camera. Lower read noise usually means better SNR and better quality of image.

Camera setting used for performance analysis is shown below:

- Full resolution
- RAW 12-bit mode
- LCG

Figure 2 shows the curves of the camera analysis data in Table 3

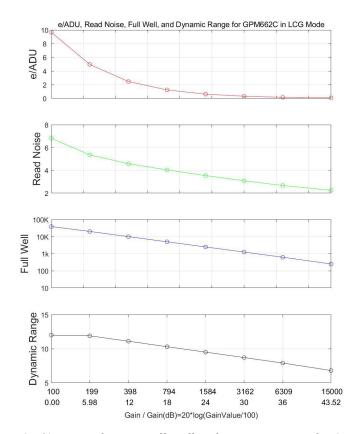


Figure 2 e/ADU, Read Noise, Full Well and Dynamic Range for GPM662C

The camera analysis data is shown in Table 3:

Table 3 Camera Analysis Data

	Sensor Analysis Data							
Gain Value	100	199	398	794	1584	3162	6309	15000
Rel Gain (dB)	0.00	5.98	12	18	24	30	36	43.52
e-/ADU	9.659	4.962	2.46	1.232	0.614	0.308	0.154	0.062
Read Noise (e-)	6.81	5.36	4.58	4.04	3.54	3.08	2.67	2.25
Full Well (ke-)	39.034	20.241	10.035	5.023	2.503	1.258	0.627	0.253
Dynamic Range (stop)	12	11.9	11.1	10.3	9.5	8.7	7.9	6.8

Camera setting used for performance analysis is shown below:

- Full resolution
- RAW 12-bit mode
- HCG

Figure 3 shows the curves of the camera analysis data in Table 4

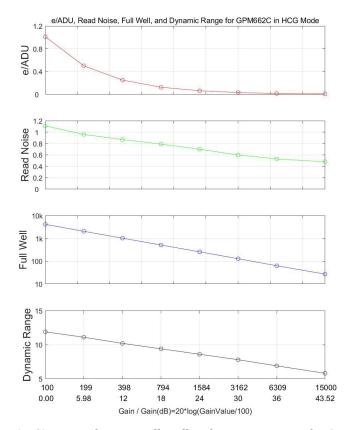


Figure 3 e/ADU, Read Noise, Full Well and Dynamic Range for GPM662C

The camera analysis data is shown in Table 4:

Table 4 Camera Analysis Data

	Sensor Analysis Data							
Gain Value	100	199	398	794	1584	3162	6309	15000
Rel Gain (dB)	0.00	5.98	12	18	24	30	36	43.52
e-/ADU	1.011	0.501	0.25	0.125	0.063	0.031	0.016	0.007
Read Noise (e-)	1.11	0.96	0.87	0.79	0.7	0.6	0.53	0.48
Full Well (ke-)	4.123	2.042	1.018	0.511	0.257	0.128	0.063	0.027
Dynamic Range (stop)	11.9	11.1	10.2	9.4	8.6	7.8	6.9	5.8

3 Product Package and Connections

3.1 Packing List



Figure 4 Packing Information of GPM662C

Table 5 GPM662C Packing List

	Standard Camera Packing List
A	Carton L:50cm W:30cm H:30cm (20pcs, 12~17Kg/ carton, 0.045m3), not shown in the photo
В	Gift box L:19.5cm W:16cm H:6cm (0.8~1.0Kg/ box)
С	One GPM series camera with a Type-CS-Mount(inner). The CS adapter ring consists of a double-layer adapter ring. The top ring can be removed to connect a 1.25-inch filter
D	High-speed USB3.0 A male to C male connectors cable /2.0m
F	1.25-inch nosepiece
G	ST4 guide cable /2.0m
Н	5mm adapter ring

3.2 Camera Dimension and Its Mount

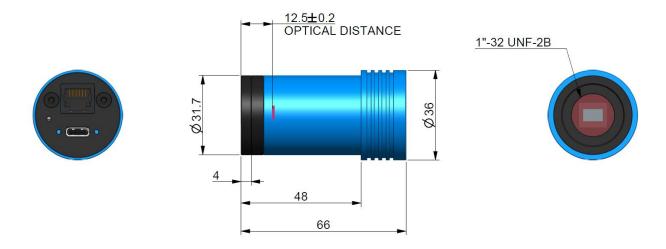


Figure 5 Dimension and Mount of GPM662C

The GPM series body, made from tough, alloy, ensures a heavy duty, workhorse solution. The camera is designed with a high-quality IR-CUT or AR to protect the camera sensor and block the IR elimination the reflection light. No moving parts included. These measures ensure a rugged, robust solution with an increased lifespan when compared to other industrial camera solutions.

Table 6 Adapter of GPM662C

Item	Specification
1	Directly with 1.25" telescope;
2	GPM + 1.25" extender with 1.25" telescope
3	Standard CS adapter connects to industrial CS-mount lens.

3.3 Camera Outline and Interface



 $Figure\ 6\ Camera\ Outline\ and\ Interface.$

Table 7 Camera Outline and Interface List

Item	Specification
1	A built-in ST4 auto guider port for the easy connection of the auto guider.
2	USB2.0 interface, Type C

4 GPM662C and Its Software

4.1 Application Installation

For software, customers are welcomed to go to our software website: https://touptek-astro.com, to download the latest ToupSky. The GPM can also be used with ASCOM, DirectShow SDK. If the third-party software is compatible with these SDK, customers can also download the software driver from our website and install the drivers into the third-party software.

ToupSky is ToupTek astronomy camera's Windows application. ToupSky is a professional software integrated with camera control, image capture & process, image browse, and analysis functions. ToupSky is born with the following features:

Windows:

- x86: XP SP3 or above; CPU supports SSE2 instruction set or above
- x64: Win7 or above

Features

- Full control of the camera
- Trigger mode and video mode support (raw format or RGB format)
- Automatic capture and quick record function
- Multi-language support
- Hardware ROI and digital binning function
- Extensive image processing functions, like image stitching, live stacking, flat field correction, dark field correction, etc.

Supported Camera:

All ToupTek astronomy cameras

4.1.1 User-friendly UI Design

- Well-arranged menus and toolbars ensure quick operating;
- The unique design of 3 sidebars -- Camera, Folders, Undo/Redo are orderly classified;
- Convenient operating method (Double click or right-click context menu) as much as possible;
- Detailed help manual;

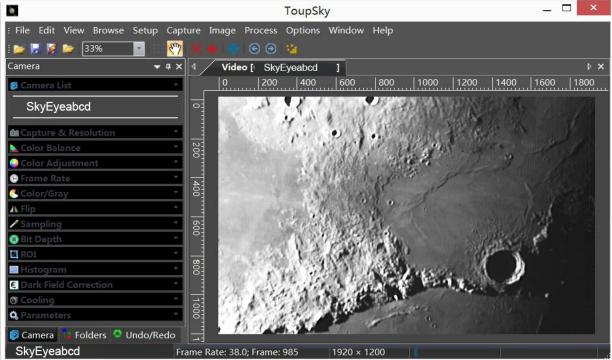


Figure 7 ToupSky and Its Video Window



Figure 8 ToupSky and Its Browse Window

4.1.2 Professional Camera Control Panel

Capture & Resolution	Set the live and snap resolution and snap the image or record video;
Exposure & Gain	Auto exposure (exposure target preset) and manual exposure (exposure time can be inputted manually); Up to 5 times gain;
Color Balance	Advanced single-click intelligent white balance setting, temperature and tint can be manually adjusted;
Color Adjustment	Hue, saturation, brightness, contrast, gamma initialization adjustment;
Frame Rate	Adjustment of frame rate available for different computer configurations;
Flip	Check the "horizontal" or "vertical" option to correct the sample direction;
Sampling	Bin mode can obtain low noise video stream; Skip mode obtains sharper and smoother video stream. Support video stream histogram extension, Negative and positive switching, Gray calibration, Clarity factor for focusing etc.
Bit Depth	Switch between 8 bits and 12 bits. 8 bits is the basic Windows image format. 12bits will have higher image quality but moderate FPS.
ROI	ROI, Region of interest. This function can set the ROI on the video window. When the ROI group was expanded, a dotted rectangle with "Handles" will appear around the video window that will let you alter the ROI. Use mouse button to adjust the ROI size. If ROI is ok, click Apply will set the video to ROI size, Defaults will return to the original size.
Dark Field Correction	To Enable the Dark Field Correction, one should capture the dark field image first. After the images are captured, the Enable button will be clickable. Checking the Enable button will enable the Dark Field correction. Unchecking it will disable the Dark Field Correction.
Cooling	Set the TE-Cooling target temperature and set the fan On/Off;
Parameters	Load, save, overwrite, import, export self-defined parameters of camera control panel (including calibration information, exposure and color setting information);

4.1.3 Practical Functions with Good Results

Video functions	Various professional functions: Video broadcast; Time lapse capture; Video record; Video stream grid; Image stitch; Video scale bar, date etc.
Image Processing and Enhancement	Control and adjust image by denoise, sharpen, color toning deinterlace, all kinds of filtering algorithm and mathematical morphology algorithm, range, binary, pseudo color, surface plot and line profile ital
Image Stacking	Image stacking adopts advanced image matching technology. With the recorded video, regardless of shifting, rotation, scaling, the high-fidelity image can be stacked to decrease the image noise.

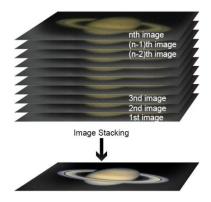


Figure 9 Image Stacking

4.1.4 Powerful Compatibility

Video Interface	Support Twain, DirectShow, SDK Package (Native C++、C#/VB.NET)
Operating System	Compatible with Microsoft® Windows® XP / Vista / 7 / 8 /10 (32 & 64 bit), Mac OSX, Linux
Language Support	Unlimited language support, currently available in Simplified Chinese, Traditional Chinese, English, Russian, German, French, Polish and Turkish

4.1.5 Hardware Requirement

	CPU: Intel Core 2 2.8GHz or higher
	Memory:2GB or more
PC Requirements	USB port: USB3.0/USB2.0 port
	Display:17" or larger
	CD-ROM

4.2 GPM662C and Dshow

ToupTekDshowAstroSetup (Click the left blue link to download)

DshowAstro is an interface driver providing Dshow standard support for ToupTek USB astronomy camera

Supported OS: Windows:

- x86: XP SP3 or above; CPU supports SSE2 instruction set or above
- x64: Win7 or above

Supported Camera:

• All ToupTek astronomy cameras

4.3 GPM662C and the 3rd Party Software

4.3.1 Support Software

No.	Software	Version	WDM	ASCOM	Native
1	PHD Guiding	2.3.0(2014)	√	√	√
2	Nebulosity	3.2.2(2014)	√	×	/
3	MaxIm DL	5.23(2013)	√	×	/
4	SharpCap	2.1(2014)	√	×	/
5	MetaGuide	5.2.0(2014)	√	/	/
6	FireCapture	2.4.05(2014)	√	/	/
7	Astroart	5.0(2014)	√	×	/

4.3.2 ASCOM Platform

All AstroCam telescope camera drivers request to install ASCOM platform, free.

http://www.ascom-standards.org/index.htm

You can download the ASCOM package from : http://ascom-standards.org/Downloads/Index.htm

4.3.3 PHD Guiding

A popular free guide software: http://openphdguiding.org/

ToupTek's telescope camera support Native/ASCOM/WDM driver to run the video.

4.3.4 Nebulosity

A popular cooled camera control/image process software directly supported via ASCOM.

4.3.5 MetaGuide

Autoguiding software with novel method to avoid the atmospheric agitation. The latest version support GCMOS01200KPB and the guide port: http://www.astrogeeks.com/Bliss/MetaGuide/

4.3.6 MAXIMDL

Famous full functional CCD Control/Image Process software. Popular used in US.

4.3.7 AstroArt

Famous full functional CCD Control/Image Process software. Popular used in Europe.

4.3.8 FireCapture

Great free planetary capture software. Support part of AstroCam series telescope camera.

4.3.9 SharpCAP

A nice free planetary capture software support WDM cameras includes AstroCam series telescope camera.

4.3.10 Registax

A popular free planetary stacking and processing software.

4.3.11 AstroStack

A planetary stacking and process software.

4.3.12 DeepSky Stacker

A free deep sky image stack and process software.

5 Service

For software upgrades, please refer to "Download" on our official website: https://touptek-astro.com/
For customers who purchase the cameras from local dealer, please contact your dealer for more inquiry.

For technical support, please contact e-mail address: <u>tpa-support@touptek.com</u>.